

# Spatial and spatio-temporal statistics for ecology

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# Section 1

## Introduction

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All ecological processes have a spatial and a spatio-temporal dimension:

- species distribution vary along the year following the season, the life stage, etc.
- ecosystem productivity is heterogeneous in space and time
- climatic variables evolve in response to climate change

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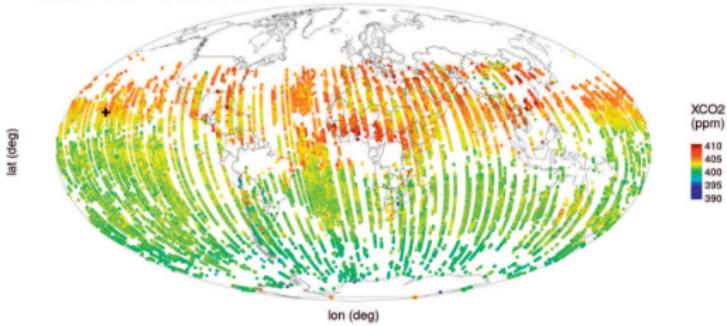
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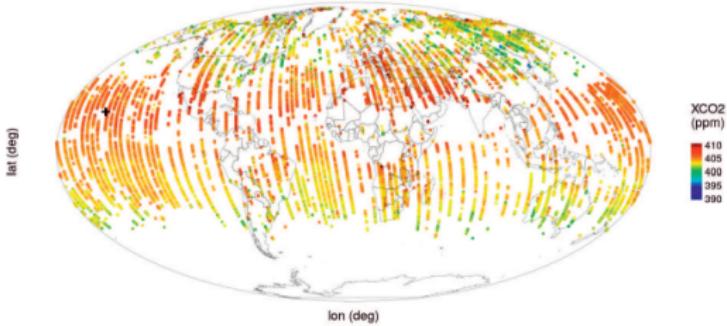
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## Examples: what are the ecological processes we would like to infer from these data?

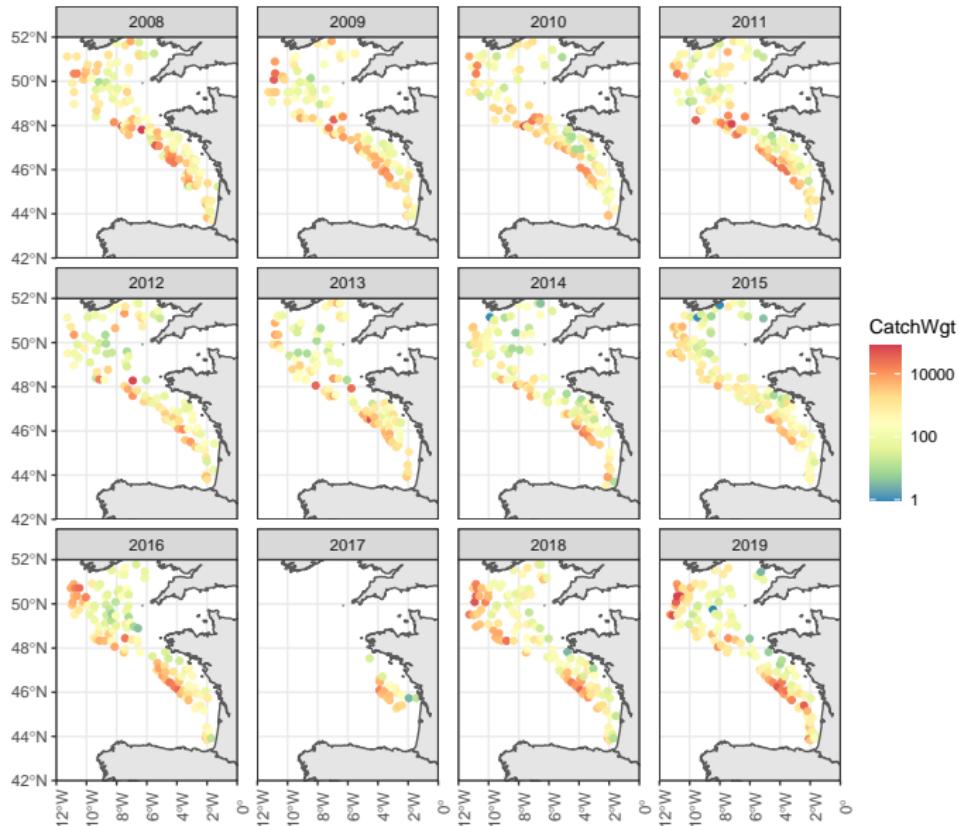
16 days of OCO-2 retrievals (V8r) 2017-01-01



16 days of OCO-2 retrievals (V8r) 2017-07-01



## *Argentina sphyraena* (EVHOE)



# Data and ecological processes

What are the characteristics of these data?

- They arise from the ecological process of interest (= they are **conditionnal** on the ecological process)
- **Noisy** (= they are not perfect observations of the ecological process)
- **Sparse** (= they do not cover the full area and some hypothesis need to be set to predict the ecological process between the sampled locations)

What are the characteristics of the ecological process we want to infer?

- **Hidden or latent**
- **Structured** (relations that structure the process?)

How to relate the data to the ecological process?

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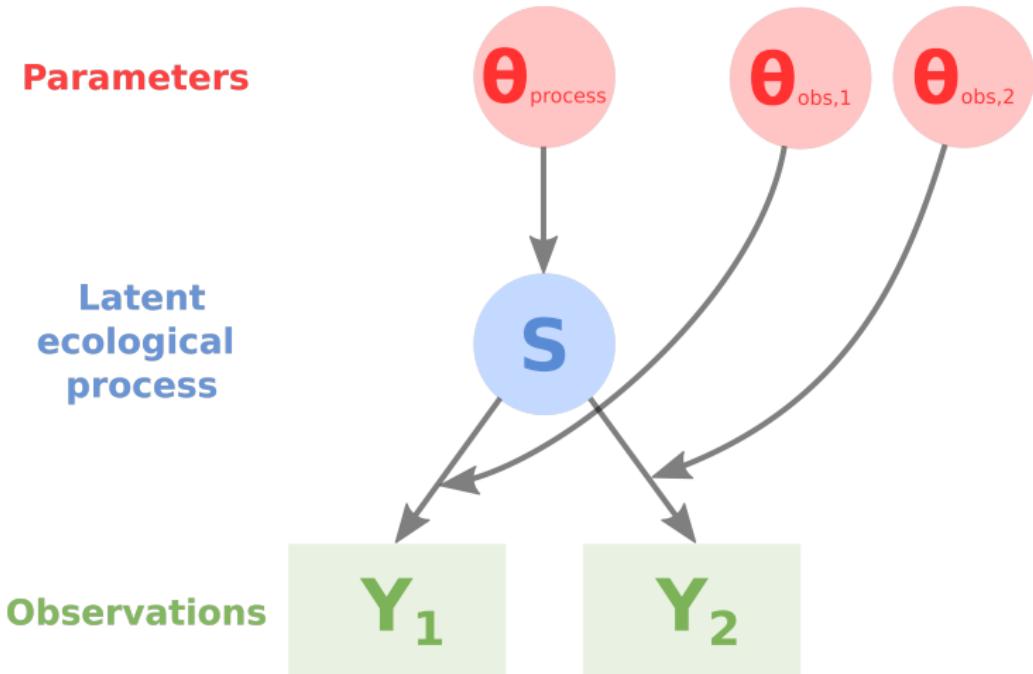
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# Hierarchical models



## Observation process

$$\mathbf{Y} | \mathbf{S}, \theta_{obs} \sim \mathcal{L}(\mathbf{S}, \theta_{obs})$$

**Y**: observations

**S**: latent field

$\mathcal{L}$ : probability distribution of **Y**

$\theta_{obs}$ : observation parameters

## Hidden process

$$\mathbf{S} | \theta_{process} \sim \mathcal{F}(\theta_{process})$$

$\mathcal{F}$ : probability distribution of **S**

$\theta_{process}$ : process parameters

- ➡ How to specify such model for space-time applications?
- ➡ What are the specificities of spatial and spatio-temporal models?

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## Section 2

### Basis of spatio-temporal statistical modeling

## Basis of spatio-temporal statistical modeling

Linear models are basis of statistical models.

For instance, in the case of EVHOE data for Argentina sphyraena, let's assume that species spatio-temporal distribution  $\mathbf{S}$  vary in space ( $x$ ), time ( $t$ ) and depends on a quadratic effect of depth.

$$\log(Y_i) \stackrel{i.i.d}{\sim} \mathcal{N}(S(x_i, t_i), \sigma^2)$$

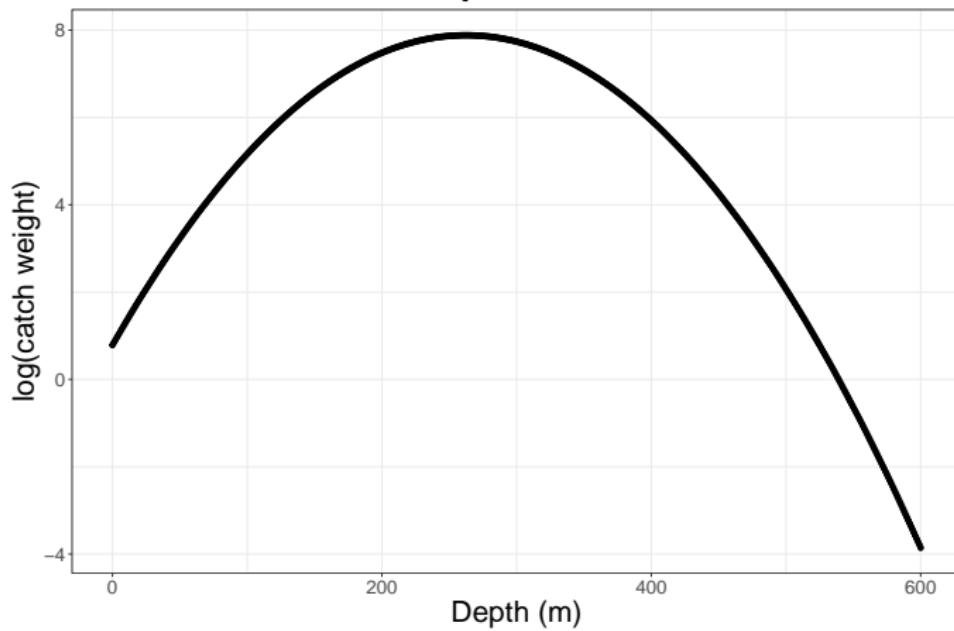
$$S(x, t) = \mu + \beta_1 \cdot \text{Depth}(x) + \beta_2 \cdot (\text{Depth}(x))^2$$

or alternatively,

$$\log(Y_i) = \mu + \beta_1 \cdot \text{Depth}(x_i) + \beta_2 \cdot (\text{Depth}(x_i)) + \epsilon_i$$

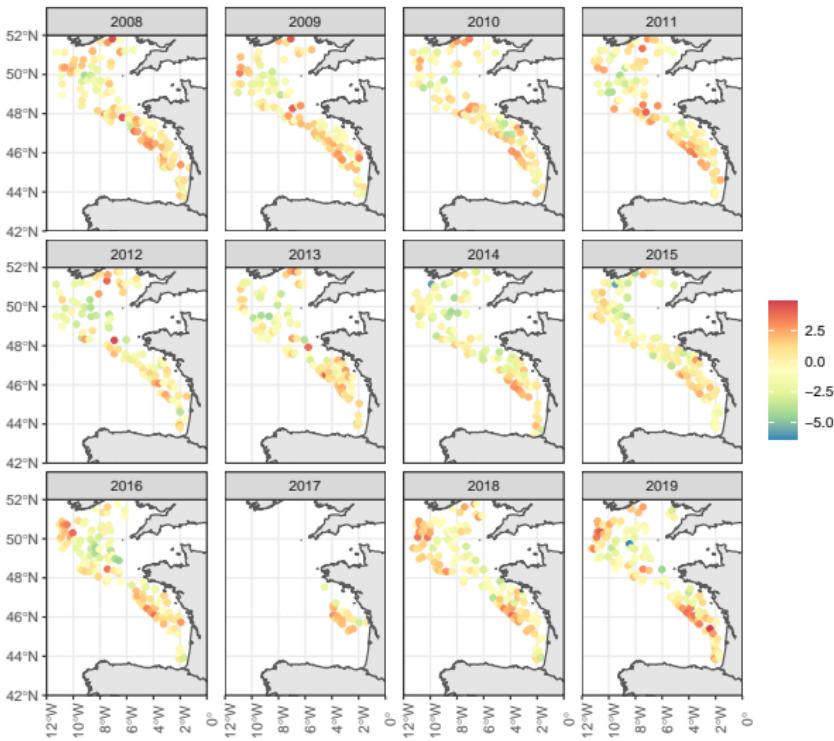
$$\epsilon_i \stackrel{i.i.d}{\sim} \mathcal{N}(0, \sigma^2)$$

## Depth effect



However, the i.i.d hypothesis is rarely satisfied in a spatio-temporal context as there can be spatio-temporal correlation structures in the residuals.

### Residuals of the linear model



## Accounting for spatial and spatio-temporal dependencies

- Introduce a Gaussian field ( $\mathcal{GF}$ )  $\delta$  in the hierarchical framework that captures spatial / spatio-temporal correlation

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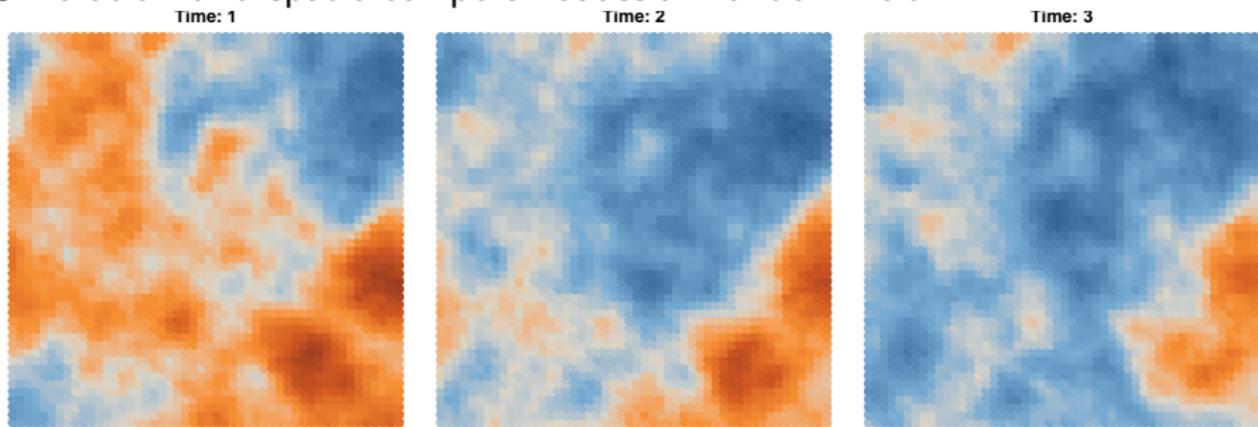
$$\delta \sim \mathcal{GF}(0, \mathcal{C}(x, y; t, r))$$

with  $\mathcal{C}(x, y; t, r)$  the spatio-temporal covariance function that controls the structure of  $\delta$

# Defining a Gaussian Field

## What is a Gaussian field?

Simulation of a spatio-temporal Gaussian random field



Probability density function of a multivariate Normal distribution:

$$f_{MN}(\delta_1, \dots, \delta_n) = (2\pi)^{-n/2} |\Sigma|^{-1/2} \exp\left(-\frac{1}{2}\delta^\top \Sigma^{-1} \delta\right)$$

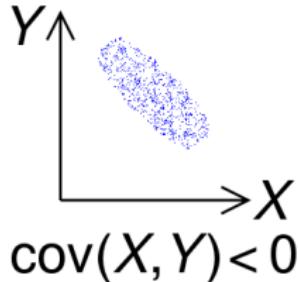
## **What is a covariance function?**

## Variance

For a random variable  $X$ :

$$\begin{aligned}\text{Var}(X) &= E[(X - E[X])^2] \\ &= E[X^2] - E[X]^2\end{aligned}$$

- Measure of dispersion. How far a set of numbers is spread out from their average value?



## Covariance

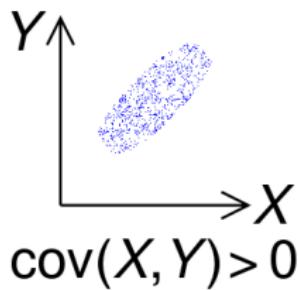
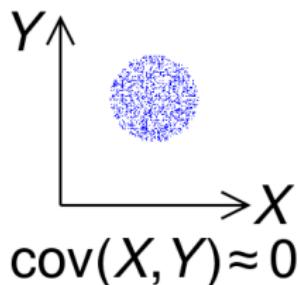
We define 2 random variables  $X$  and  $Y$ .

$$\text{Cov}(X, Y) \equiv E[(X - E[X])(Y - E[Y])]$$

- Measure of the joint variability of two random variables.

The **correlation coefficient** is a normalized measure of the covariance.

$$\rho_{X,Y} = \frac{\text{cov}(X, Y)}{\sigma_X \sigma_Y}$$

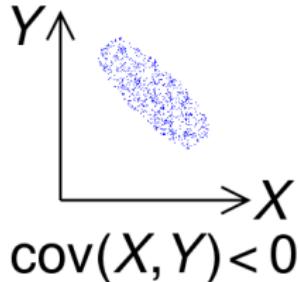


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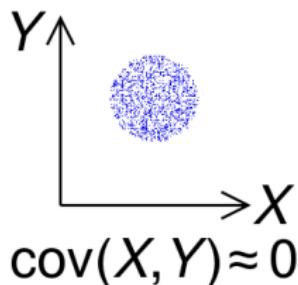


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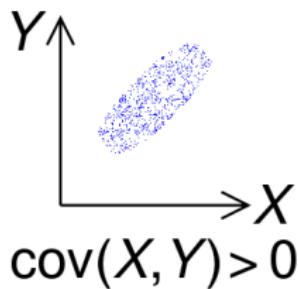
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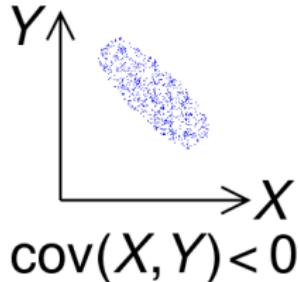


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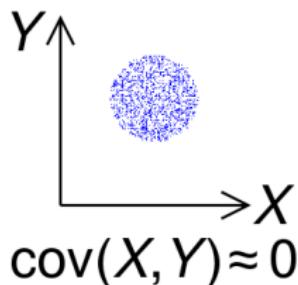


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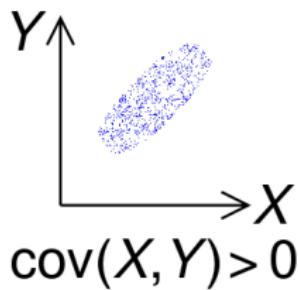
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## Variance-covariance matrix

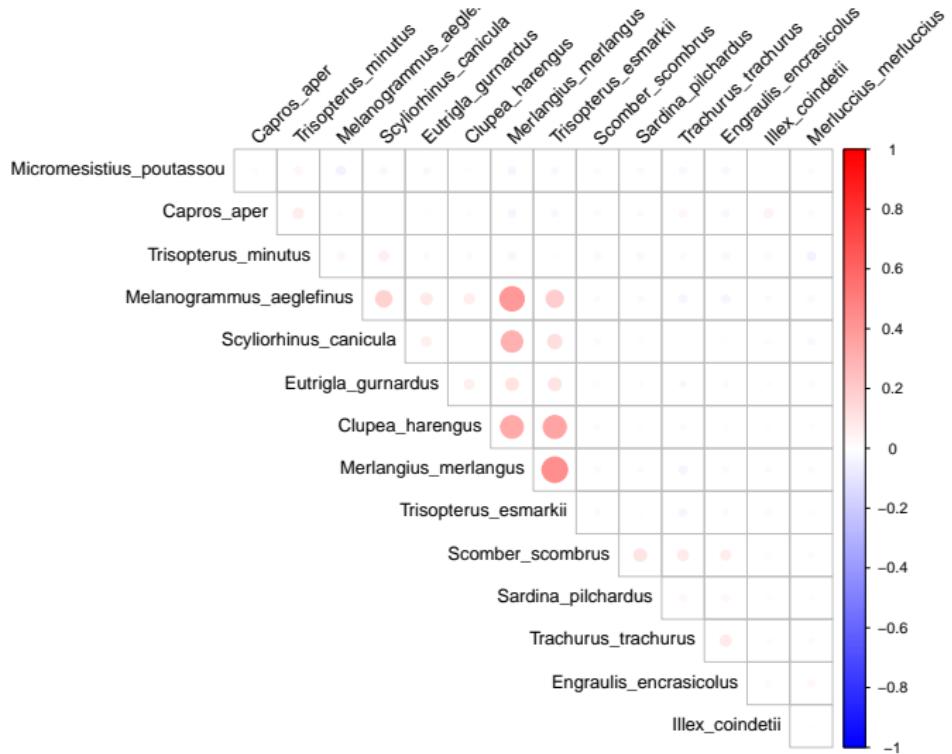
For a vector of  $p$  random variables  $\mathbf{X} = (X_1, \dots, X_p)$ :

$$\text{Var}(\mathbf{X}) = \begin{pmatrix} \text{Var}(X_1) & \text{Cov}(X_1, X_2) & \cdots & \text{Cov}(X_1, X_p) \\ \text{Cov}(X_2, X_1) & \ddots & \cdots & \vdots \\ \vdots & \vdots & \ddots & \vdots \\ \text{Cov}(X_p, X_1) & \cdots & \cdots & \text{Var}(X_p) \end{pmatrix}$$

The correlation matrix is the normalized version of the covariance matrix.

$\text{Var}(\mathbf{X})^{-1}$  is the precision matrix.

# Correlation matrix for the catch weights of the 15 most abundant species of EVHOE data



## Spatial variance-covariance

Let's consider the random effect  $\delta(x, t)$  at a given time step  $t$  that we rewrite  $\delta(x)$ .

The spatial variance-covariance matrix of  $\delta(x)$  can be written:

$$\text{Var}(\delta(.)) = \begin{pmatrix} \text{Var}(\delta(x_1)) & \text{Cov}(\delta(x_1), \delta(x_2)) & \cdots & \text{Cov}(\delta(x_1), \delta(x_n)) \\ \text{Cov}(\delta(x_2), \delta(x_1)) & \ddots & \cdots & \vdots \\ \vdots & \vdots & \ddots & \vdots \\ \text{Cov}(\delta(x_n), \delta(x_1)) & \cdots & \cdots & \text{Var}(\delta(x_n)) \end{pmatrix}$$

► **No spatial representation of the covariance matrix per se**

The spatial dependencies between the values of  $\delta(.)$  is modeled through a variance-covariance function  $\mathcal{C}(x, y)$  (spatial only here).

Some base properties are required to define a spatial covariance function.

## Necessary properties for spatial covariance matrix

### Non-negative and positive-definiteness

A function  $\{\mathcal{C}(x, y) : x, y \in D\}$  defined on  $D$  is said to be **non-negative-definite**, if for any complex numbers  $\{a_i : i = 1, \dots, m\}$ , any  $\{x_i : i = 1, \dots, m\}$  in  $D$ , and any integer  $m$ , we have

$$\sum_{i=1}^m \sum_{j=1}^m a_i \bar{a}_j \mathcal{C}(x_i, x_j) \geq 0$$

where  $\bar{a}$  denotes the complex conjugate of  $a$ .

To be *valid*, a covariance function must be non-negative definite.

A function is **positive-definite** when the inequality below is strictly positive whenever  $(a_1, \dots, a_m)'$  is a nonzero vector.

## Additionnal (convenient) properties for spatial covariance matrix

### Stationarity

There are 2 main kinds of stationarity: strong and second-order (or weak) stationarity.

- $\delta(x)$  is strongly stationary when the two probability measures defining  $\delta(x)$  and  $\delta(x + h)$  are equivalent for all  $h \in \mathbb{R}^d$ .
- $\delta(x)$  is second-order (or weakly) stationary when it has a constant expectation and a stationary covariance function.

### Isotropy

Spatial isotropy corresponds to  $\text{Cov}(\delta(x), \delta(y)) \equiv \mathcal{C}(\|x - y\|)$

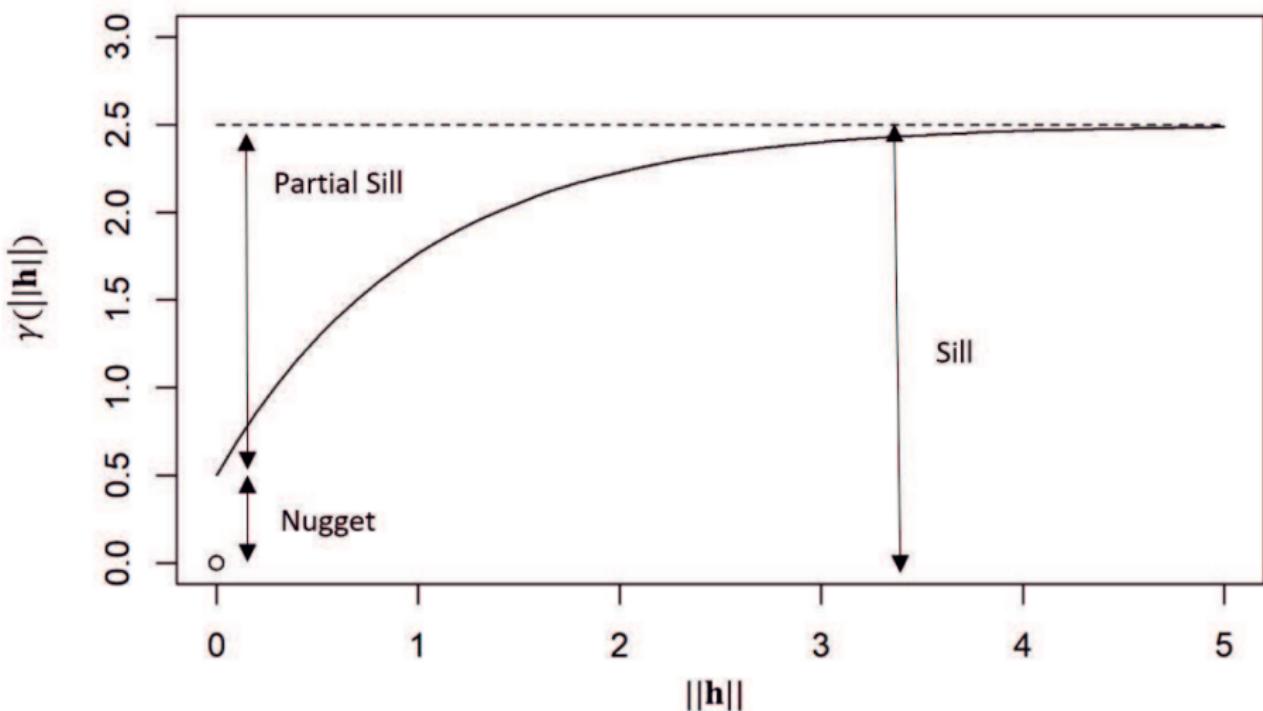
### Intrinsic

$\delta(x)$  is intrinsic if for all  $(x) \in \mathbb{R}^d$ ,  $h \in \mathbb{R}$ ,  $\delta(x + h) + \delta(x)$  is second-ordered stationary (constant expectation and stationary covariance).

## Some examples of isotropic covariance functions

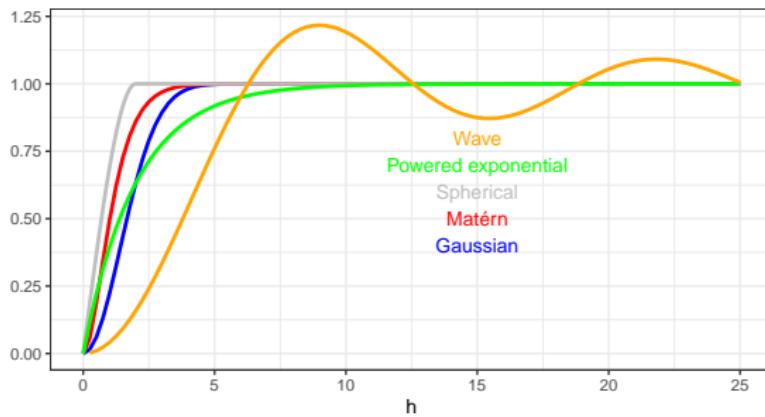
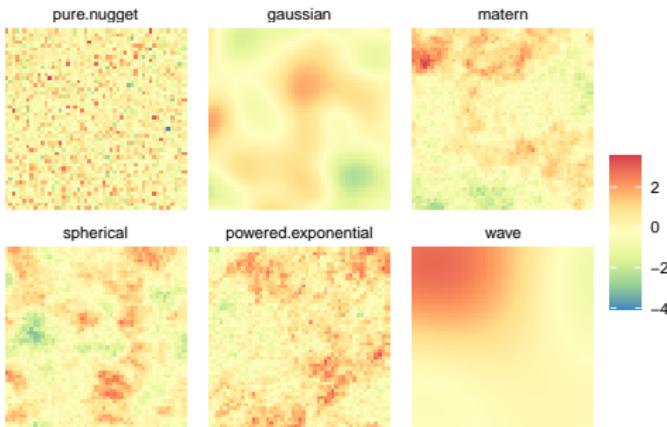
Model	Covariance function, $C(d)$
Linear	$C(d)$ does not exist
Spherical	$C(d) = \begin{cases} 0 & \text{if } d \geq 1/\phi \\ \sigma^2 \left[ 1 - \frac{3}{2}\phi d + \frac{1}{2}(\phi d)^3 \right] & \text{if } 0 < d \leq 1/\phi \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Exponential	$C(d) = \begin{cases} \sigma^2 \exp(-\phi d) & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Powered exponential	$C(d) = \begin{cases} \sigma^2 \exp(- \phi d ^p) & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Gaussian	$C(d) = \begin{cases} \sigma^2 \exp(-\phi^2 d^2) & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Rational quadratic	$C(d) = \begin{cases} \sigma^2 \left( 1 - \frac{d^2}{(\phi+d^2)} \right) & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Wave	$C(d) = \begin{cases} \sigma^2 \frac{\sin(\phi d)}{\phi d} & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Power law	$C(d)$ does not exist
Matérn	$C(d) = \begin{cases} \frac{\sigma^2}{2^{\nu-1}\Gamma(\nu)} (\phi d)^{\nu} K_{\nu}(\phi d) & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$
Matérn at $\nu = 3/2$	$C(d) = \begin{cases} \sigma^2 (1 + \phi d) \exp(-\phi d) & \text{if } d > 0 \\ \tau^2 + \sigma^2 & \\ d = 0 & \end{cases}$

Another (more general) way to define spatial structure is to define the **covariogram**  $\text{Var}(Y(x) - Y(y)) = 2\gamma(x, y)$  (with  $\gamma()$  the **semivariogram**) and its stationary version  $2\gamma(h) = 2(\gamma' - \langle \rangle); h \in \mathbb{R}$ .



## Some examples of isotropic variograms

Model	Variogram, $\gamma(d)$
Linear	$\gamma(d) = \begin{cases} \tau^2 + \sigma^2 d & \text{if } d > 0 \\ 0 & \text{if } d = 0 \end{cases}$
Spherical	$\gamma(d) = \begin{cases} \tau^2 + \sigma^2 & \text{if } d \geq 1/\phi \\ \tau^2 + \sigma^2 [\frac{3}{2}\phi d - \frac{1}{2}(\phi d)^3] & \text{if } 0 < d \leq 1/\phi \\ 0 & \text{if } d = 0 \end{cases}$
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Rational quadratic	$\gamma(d) = \begin{cases} \tau^2 + \frac{\sigma^2 d^2}{(\phi + d^2)} & \text{if } d > 0 \\ 0 & \text{if } d = 0 \end{cases}$
Wave	$\gamma(d) = \begin{cases} \tau^2 + \sigma^2(1 - \frac{\sin(\phi d)}{\phi d}) & \text{if } d > 0 \\ 0 & \text{if } d = 0 \end{cases}$
Power law	$\gamma(d) = \begin{cases} \tau^2 + \sigma^2 d^\lambda & \text{if } d > 0 \\ 0 & \text{if } d = 0 \end{cases}$
Matérn	$\gamma(d) = \begin{cases} \tau^2 + \sigma^2 \left[1 - \frac{(\phi d)^\nu}{2^{\nu-1} \Gamma(\nu)} K_\nu(\phi d)\right] & \text{if } d > 0 \\ 0 & \text{if } d = 0 \end{cases}$
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# Moving to spatio-temporal

The ideas are similar, but we add temporal correlations in the expression of the random effect  $\delta$ .

For instance, let's introduce the model:

$$\delta(x, t) = \varphi \cdot \delta(x, t - 1) + \omega(x, t) \text{ for } t = 2, \dots, T$$

- $\varphi \in ]-1; 1[$  is the autoregressive temporal term
- $\omega(x, t)$  is a purely spatial GRF
- $\omega(x, 1)$  derives from the stationary distribution  $\mathcal{N}(0, \sigma^2 / (1 - \varphi^2))$

$$\mathcal{C}(\omega(x, t), \omega(y, r)) = \begin{cases} 0 & \text{if } t \neq r \\ \sigma_\omega^2 \cdot \text{Cor}(h) & \text{if } t = r \end{cases} \quad \text{for } x \neq y$$

where  $h = \|x - y\| \in \mathbb{R}$  and  $\text{Cor}(h)$  is the Matérn correlation function.

The variance-covariance matrix can be written as:

$$\mathcal{C}(\delta(x, t), \delta(y, r)) = \frac{\varphi^{|t-r|}}{1 - \varphi^2} \sigma_\omega^2 \text{Cor}(h)$$

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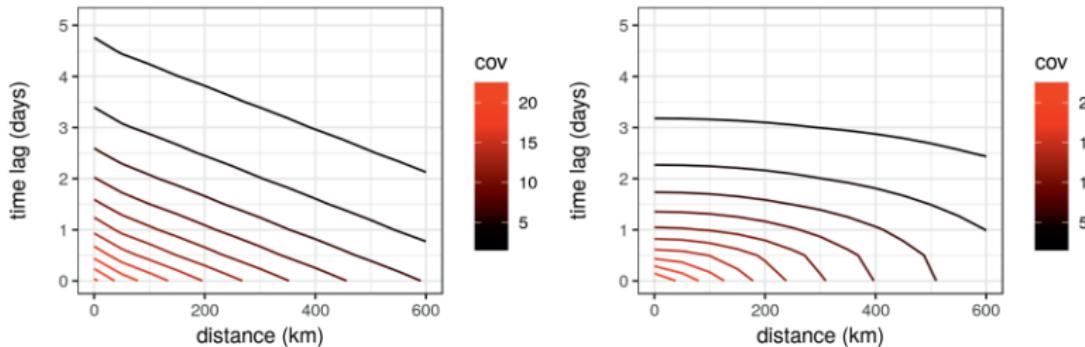
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Such kind of process is:

- stationary:  $\mathcal{C}(x, y; t, r) = \mathcal{C}(h, \tau)$  with  $h = ||x - y||$  and  $\tau = |t - r|$ .
- separable:  $\mathcal{C}(h, \tau) = \mathcal{C}^{(x)}(h) \cdot \mathcal{C}^{(t)}(\tau)$
- fully symmetric:  $\text{cov}(\delta(x, t); \delta(y, r)) = \text{cov}(\delta(y, t); \delta(x, r))$

► But other non-separable covariance function exist and might be desirable to model complex interactions between spatial and temporal correlations.

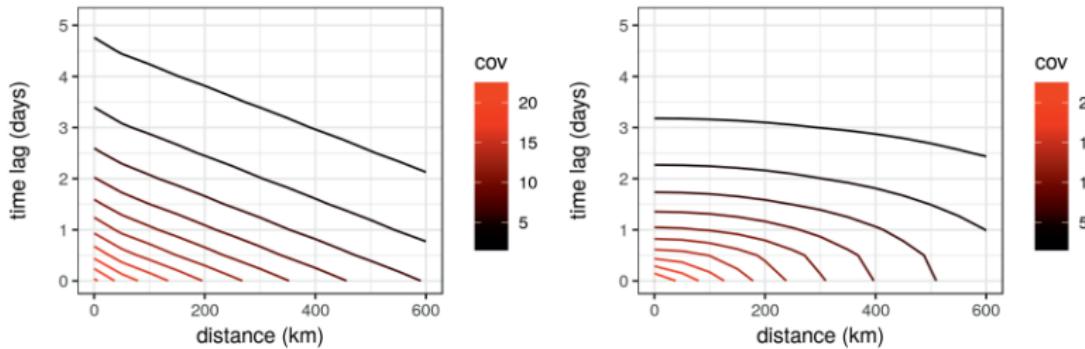


Separable covariance function (left) vs. non-separable covariance function (right)

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Separable covariance function (left) vs. non-separable covariance function (right)

## Section 3

### Inference method

# Kriging

Basis of geostatistics ➔ Minimum mean-squared error approach

Let's define observations  $\mathbf{Y} = (Y_1, Y_2, \dots, Y_n)$

A linear predictor of  $Y_0$  based on  $\mathbf{Y}$  would take the form  $\sum \ell_i Y_i + \alpha_0$

Using squared error loss, the best linear prediction would minimize

$$E \left[ Y_0 - \left( \sum \ell_i Y_i + \alpha_0 \right) \right]^2 \text{ over } \ell_i \text{ and } \alpha_0$$

Setting  $\sum \ell_i = 1$ ,  $\alpha_0 = 0$ ,  $a_0 = 1$  and  $a_i = -\ell_i$ , the criterion becomes  $E[\sum_{i=0}^n a_i Y_i]^2$  with  $\sum a_i = 0$ .

$$\text{And } E[\sum_{i=0}^n a_i Y_i]^2 = - \sum_i \sum_j a_i a_j \gamma(||x_i - x_j||)$$

The  $\ell_i$  are then estimated by solving constrained optimization on this criterion

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## Maximum likelihood approach

Maximum likelihood methods consist in looking for the parameter values that maximize the likelihood

$$P(\mathbf{Y}|\boldsymbol{\theta})$$

For hierarchical models, this suppose to make an integral over the random effect  $\delta$  (with  $q$  the size of  $\delta$ ).

$$P(\mathbf{Y}|\boldsymbol{\theta}) = \int_{\mathbb{R}^q} P(\mathbf{Y}, \boldsymbol{\delta}|\boldsymbol{\theta}) d\boldsymbol{\delta}$$

➡  $q$  is often very high in spatio-temporal model

To bypass the computational burden:

- Laplace approximation
- Pseudo-likelihood

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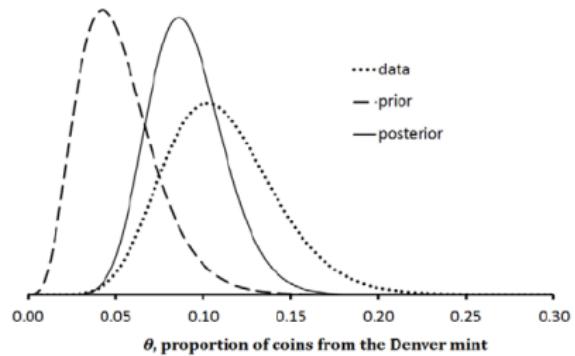
...

# Bayesian inference

Bayesian Paradigm:

$$P(\theta, \delta | \mathbf{Y}) = \frac{P(\mathbf{Y} | \theta, \delta) \cdot P(\theta, \delta)}{P(\mathbf{Y})}$$

- $P(\theta, \delta | \mathbf{Y})$ : posterior distribution of parameters and random effect
- $P(\mathbf{Y} | \theta, \delta)$ : likelihood of the data
- $P(\theta, \delta)$ : prior distribution of parameters and random effect
- $P(\mathbf{Y})$ : marginal likelihood of the data (= normalizing constant)



## MCMC: Random sampling in the joint posterior distribution $P(\delta, \theta | \mathbf{Y})$ (Metropolis-Hastings algorithm)

$$P(\delta, \theta | \mathbf{Y}) \propto P(\theta) \cdot P(\delta | \theta) \cdot P(\mathbf{Y} | \delta, \theta)$$

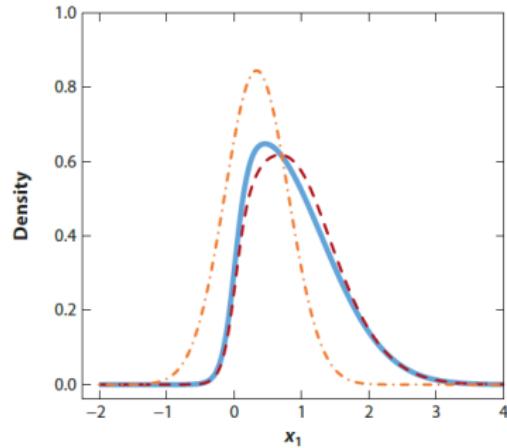
→ In practice, the estimation can be very long due to the size of  $\delta$

⇒ INLA: instead of estimating the joint distribution  $P(\delta, \theta | \mathbf{Y})$ , approximate the marginal distribution  $P(\theta | \mathbf{Y})$  and  $P(\delta_i | \theta, \mathbf{Y})$ .

$$\tilde{P}(\theta | \mathbf{y}) \propto \frac{P(\delta, \theta, \mathbf{Y})}{\tilde{P}_{\mathbf{G}}(\delta | \theta, \mathbf{Y})} \Big|_{\delta=\delta^*(\theta)}$$

$$\tilde{P}(\theta_j | \mathbf{Y}) = \int \tilde{P}(\theta | \mathbf{Y}) d\theta_{-j}$$

$$\tilde{P}(\delta_i | \mathbf{y}) = \int \tilde{P}(\delta_i | \theta, \mathbf{Y}) \tilde{P}(\theta | \mathbf{Y}) d\theta$$



Marginal distribution of  $\delta_1$  (blue), Laplace approximation (red) and Gaussian approximation (orange).

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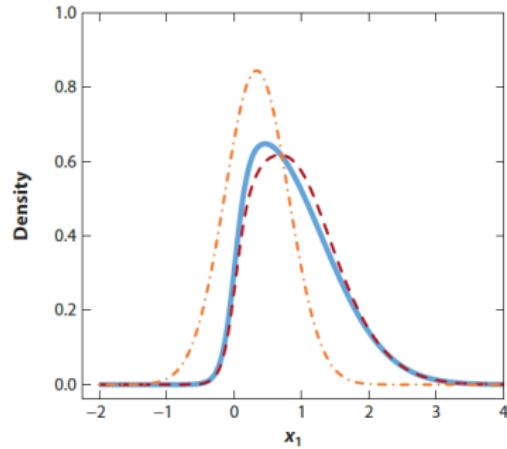
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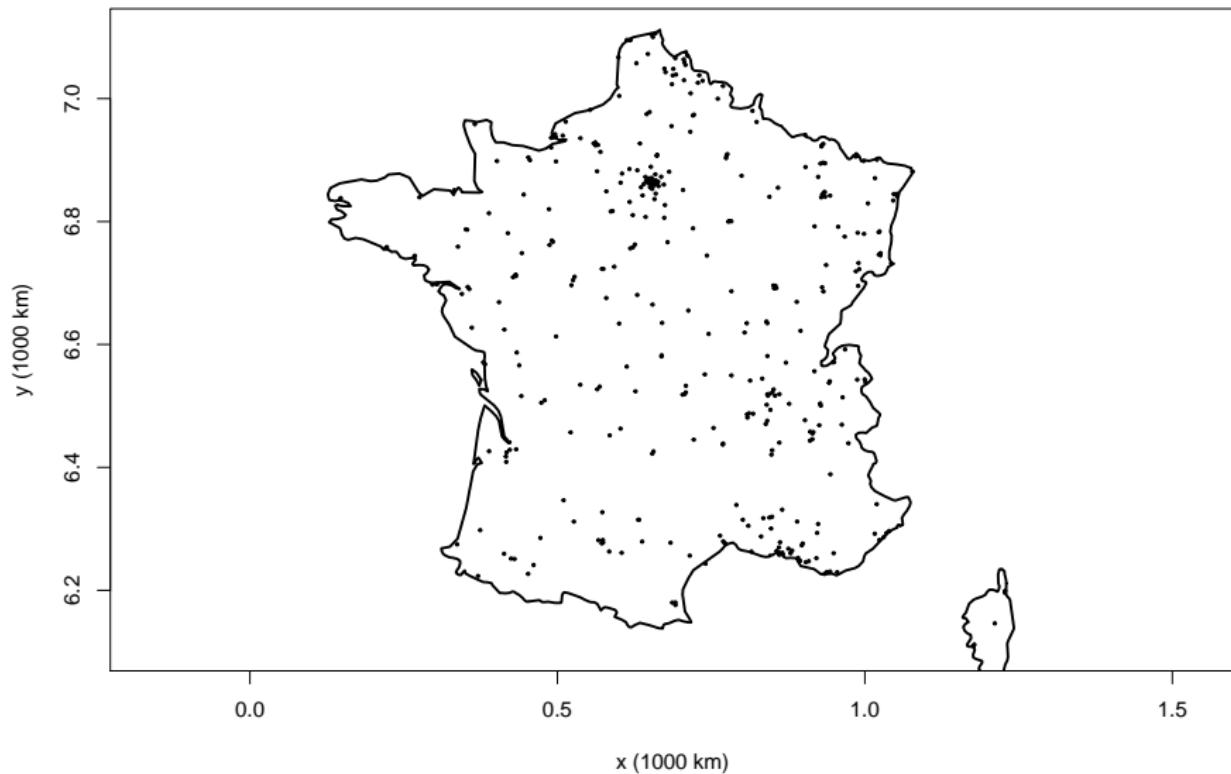


Marginal distribution of  $\delta_1$  (blue), Laplace approximation (red) and Gaussian approximation (orange).

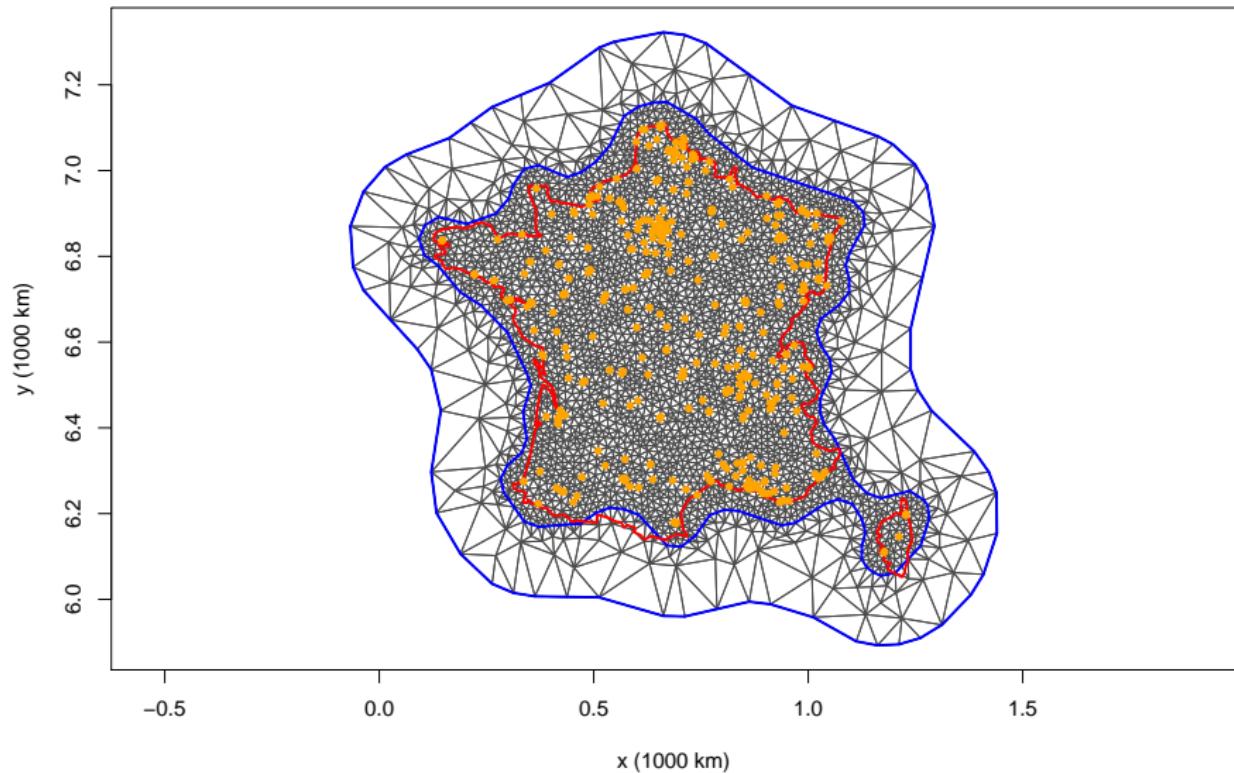
## Section 4

Some practice!

# Data



## Creating the mesh



# Building the SPDE model

```
myspde=inla.spde2.pcmatern(mesh=mesh,
                             alpha=2,
                             prior.range=c(.01,.1),
                             prior.sigma=c(25,0.5))

# Observation matrix to make the connection
# between mesh nodes and observation sites:
A=inla.spde.make.A(mesh,loc=xy_OBS)

# Vector of indices for the nodes of the mesh
idx.spatial=inla.spde.make.index("spatial",
                                  n.spde=mesh$n)
```

## Building the regression part

First, we create covariate dataframe: we add the intercept term 1, since in R-INLA it's more comfortable to include the intercept explicitly.

```
covar.df=data.frame(intercept=1,  
                     x=xy_OBS[,1],  
                     y=xy_OBS[,2])  
  
# Create the stack with all the data and indices:  
mystack=inla.stack(data=list(pm10=OBS_daily2014$PM10),  
                     A=list(A,1),  
                     effects=list(idx.spatial,covar.df))  
  
# We write the formula for the model:  
myformula=pm10~-1+intercept+f(spatial,model=mypde)
```

Note: since we handle the intercept explicitly, we have to include “-1”). Recall that we have given the name “spatial” to the index of the spatial effect.

# Fitting the model with R-INLA

The following `inla` run should take approximatively 1 minute.

```
fit=inla(myformula,
          data=inla.stack.data(mystack),
          family="gaussian",
          control.predictor=list(A=inla.stack.A(mystack),
                                 compute=FALSE),
          control.inla=list(int.strategy="eb",
                            strategy="gaussian"),
          verbose=F)
```

```

## 
## Call:
##   c("inla(formula = myformula, family = \"gaussian\", data =
##     inla.stack.data(mystack), ", " verbose = F, control.predictor = list(A
##       = inla.stack.A(mystack), ", " compute = FALSE), control.inla =
##       list(int.strategy = \"eb\", ", " strategy = \"gaussian\"))")
## Time used:
##   Pre = 2.12, Running = 106, Post = 0.506, Total = 109
## Fixed effects:
##           mean      sd 0.025quant 0.5quant 0.975quant    mode kld
## intercept 18.039 0.323      17.404   18.039      18.674 18.039    0
##
## Random effects:
##   Name      Model
##   spatial  SPDE2 model
##
## Model hyperparameters:
##           mean      sd 0.025quant 0.5quant
## Precision for the Gaussian observations 0.009 0.000      0.009 0.009
## Range for spatial                      0.022 0.005      0.013 0.022
## Stdev for spatial                      4.248 0.358      3.659 4.205
##           0.975quant    mode
## Precision for the Gaussian observations      0.009 0.009
## Range for spatial                        0.032 0.022
## Stdev for spatial                       5.054 4.085
##
## Expected number of effective parameters(stdev): 196.69(0.00)
## Number of equivalent replicates : 381.75
##
## Marginal log-Likelihood: -283189.38

```

# Predictions with R-INLA

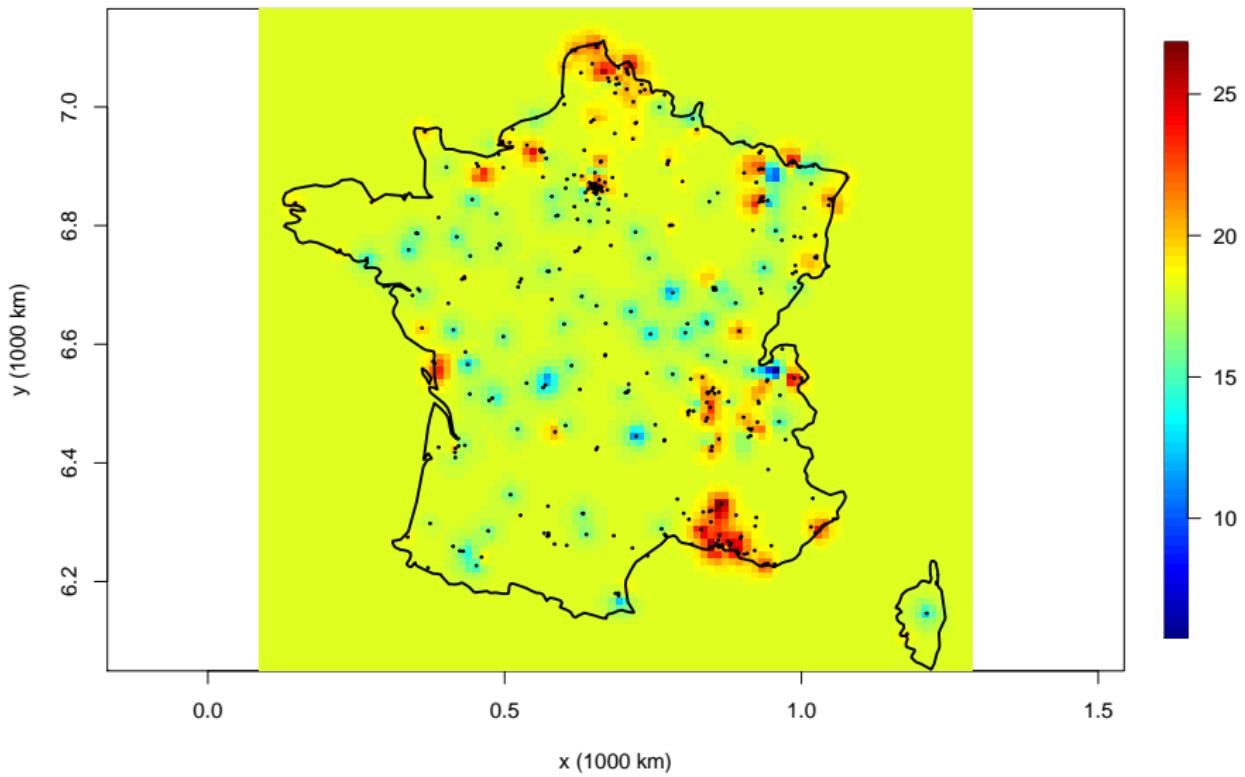
To build predictions, the input data (`inla.stack`) must be slightly modified. NA values are added to the vector of observations and `inla` will predict latent field values at these locations.

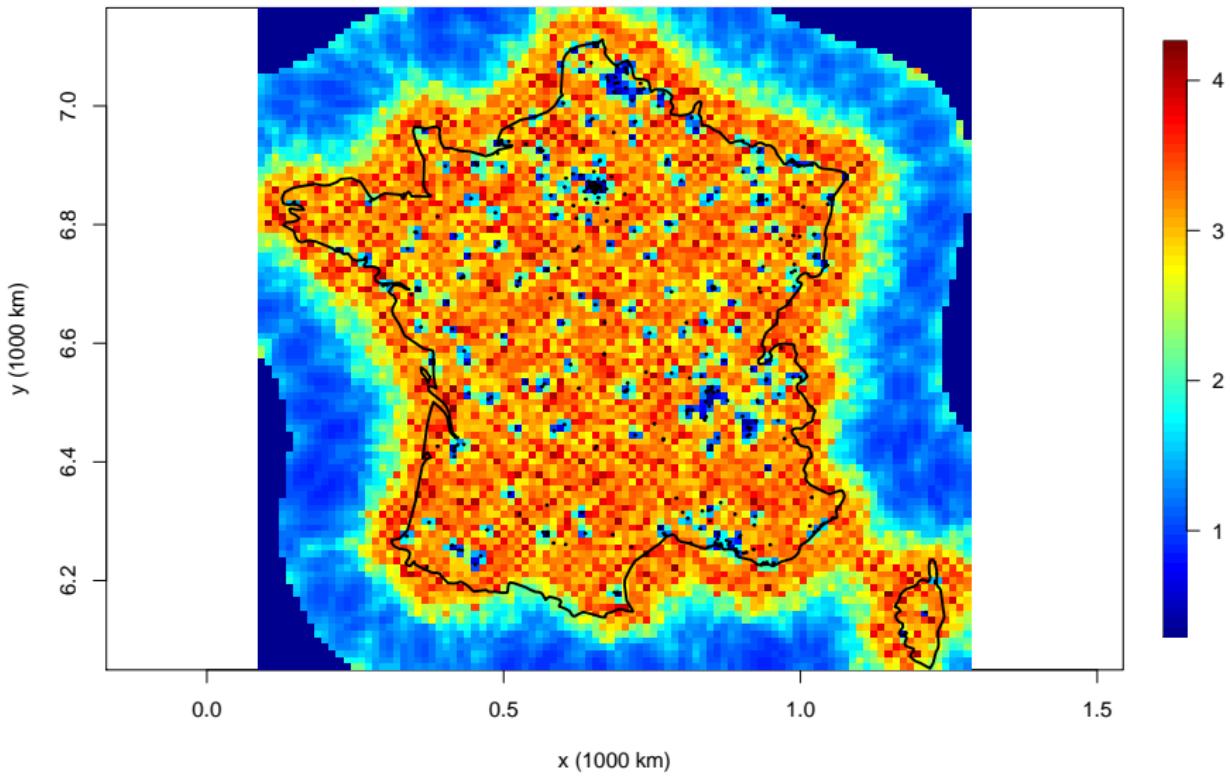
```
# Observation matrix to make the connection between mesh nodes and observation sites:  
proj_grid=inla.mesh.projector(mesh,  
                               xlim=range(bound$loc[,1]),  
                               ylim=range(bound$loc[,2]),  
                               dims=c(100,100))  
  
# Grid on which we calculate the predictions:  
xygrid=as.matrix(expand.grid(proj_grid$x,proj_grid$y))  
  
# Put prediction coordinates and observation coordinates into A:  
A=inla.spde.make.A(mesh,  
                    loc=rbind(xygrid,xy_OBS))  
  
# We add the intercept term 1, since in RINLA it's more comfortable to include the intercept explicitly.  
covar=df=data.frame(intercept=1,  
                     x=c(xygrid[,1],  
                           xy_OBS[,1]),  
                     y=c(xygrid[,2],  
                           xy_OBS[,2]))  
  
# Create the stack with all the data and indices: add NA data here  
mystack=inla.stack(data=list(pm10=c(rep(NA,nrow(xygrid)),  
                                         OBS_daily2014$PM10)),  
                     A=list(A,1),  
                     effects=list(idx.spatial,covar=df),  
                     tag="mytag")
```

When fitting, we now set `compute=TRUE` (to calculate fitted and predicted values) and `link=1` (although only necessary when using non-identity link).

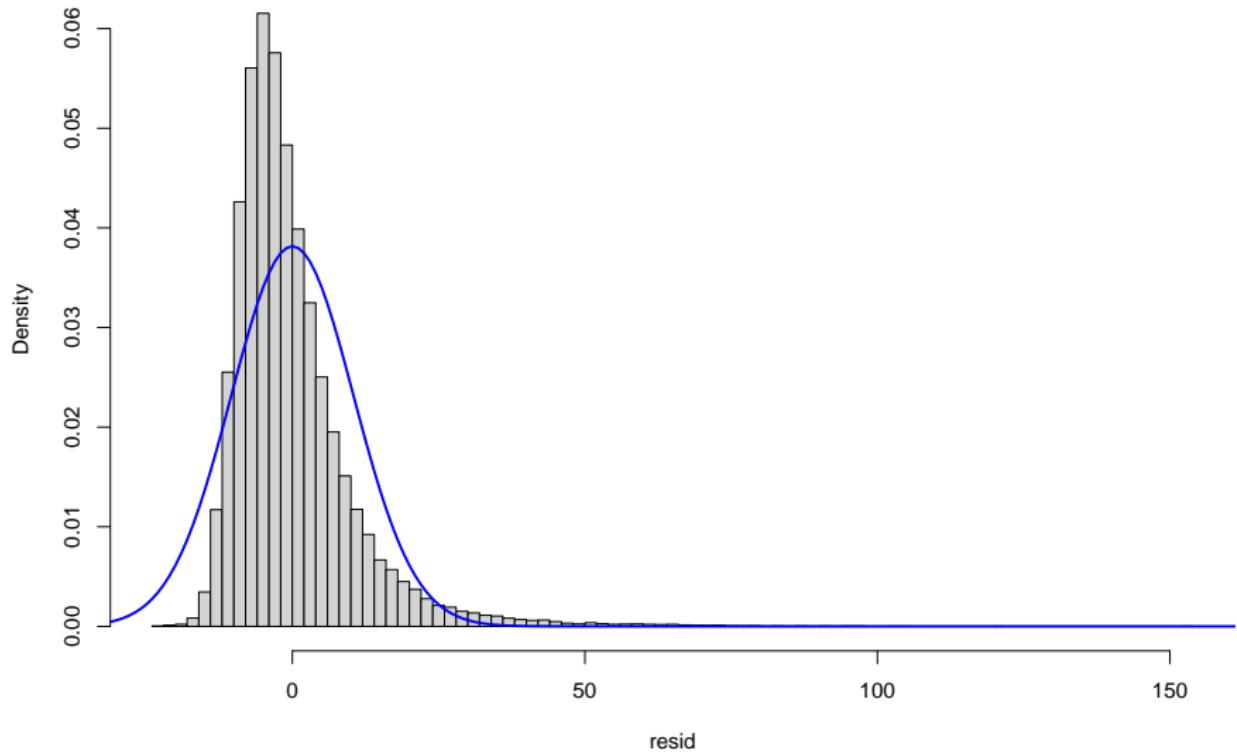
Running INLA can now take considerably longer since we calculate the posterior estimation `y.hat` for each observation point `y` (`compute=TRUE`).

```
fit=inla(myformula,
          data=inla.stack.data(mystack),
          family="gaussian",
          control.predictor=list(A=inla.stack.A(mystack),
                                 compute=TRUE,
                                 link=1),
          control.inla=list(int.strategy="eb",
                            strategy="gaussian"),
          verbose=TRUE)
```



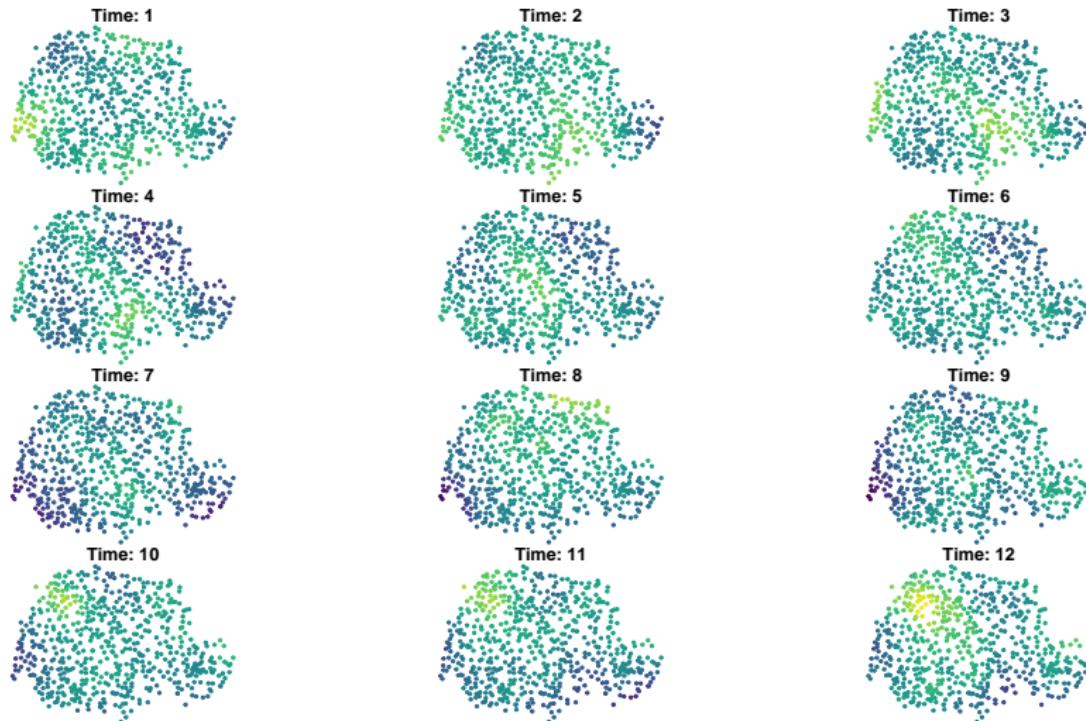


### Histogram of resid

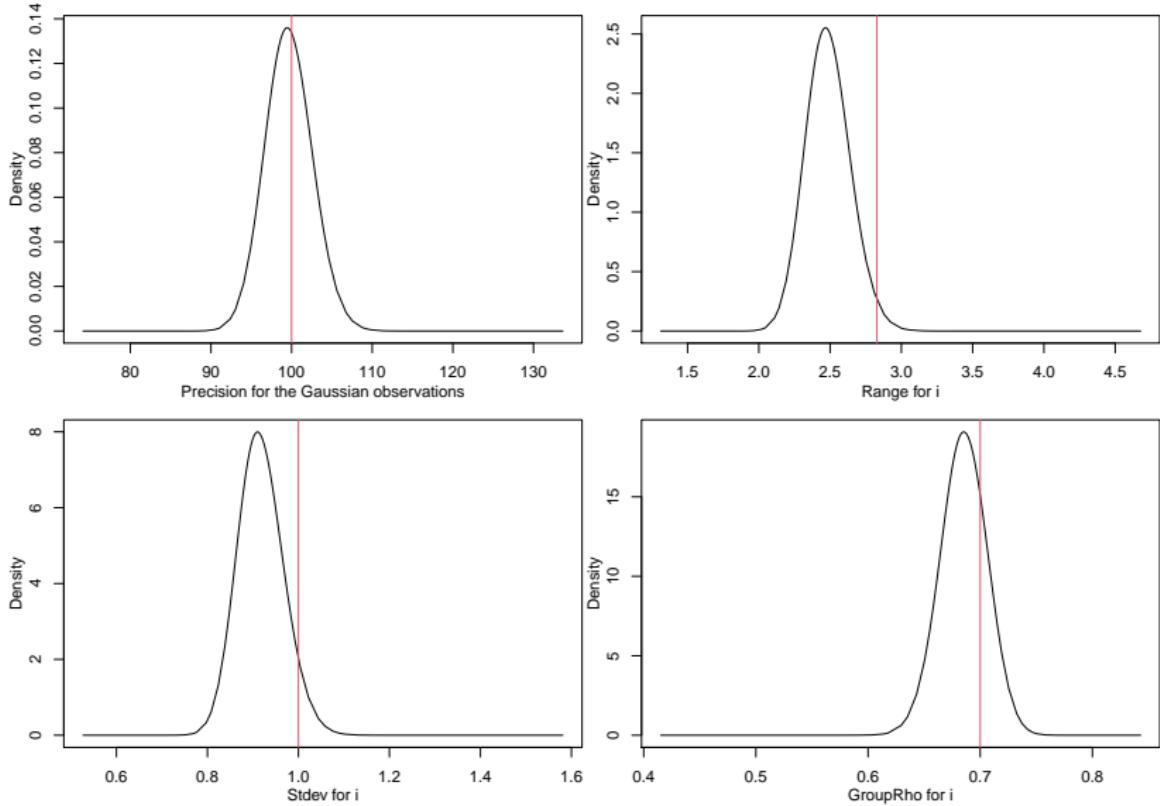


# Spatio-temporal application

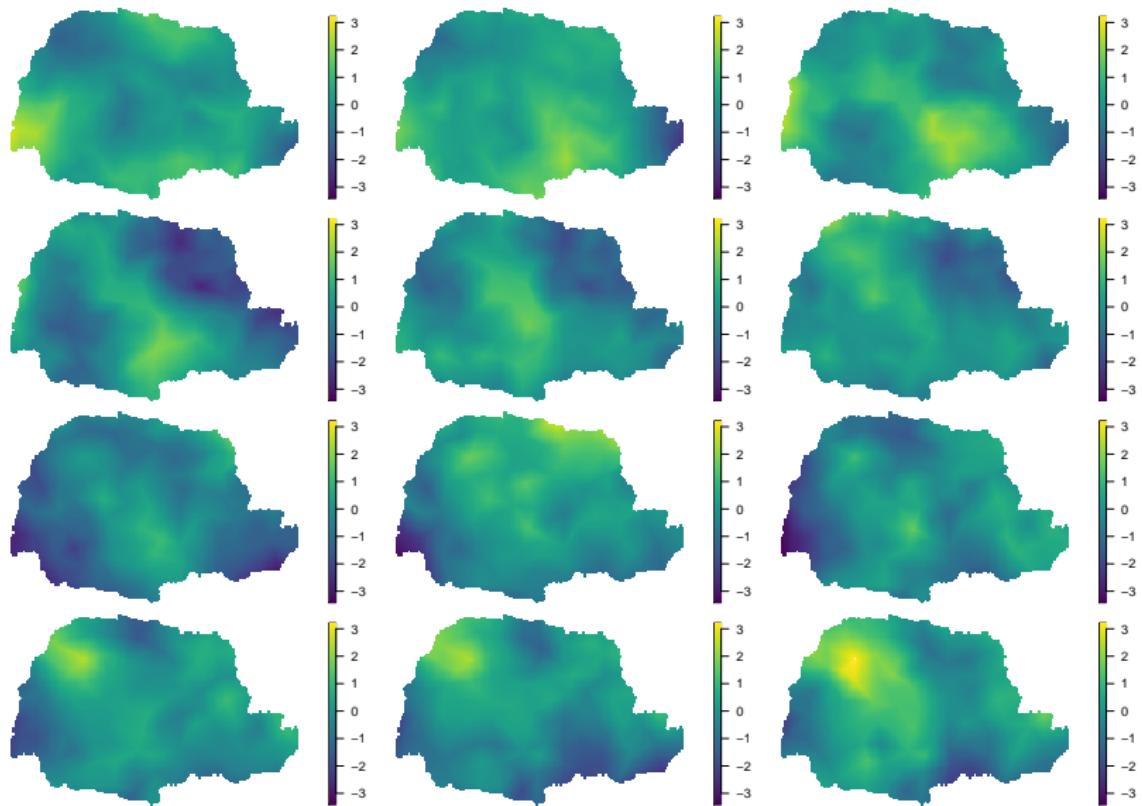
Simulating and preparing the data



## Fitting the model - model estimates



## Fitting the model - model spatial predictions

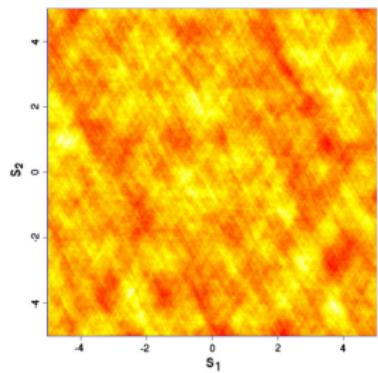


## Section 5

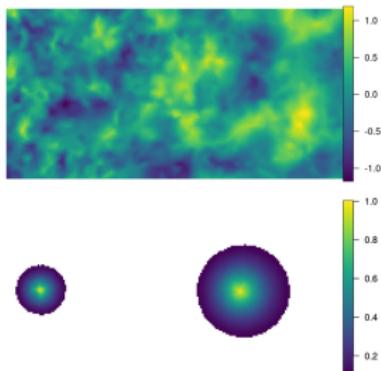
Few words on some subject that we did not address  
here

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## Anisotropy

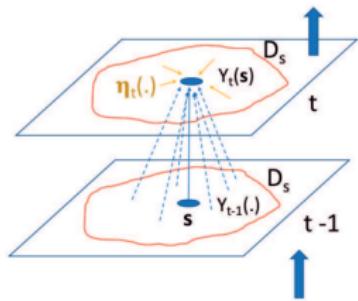


## Non-stationarity



## Dynamic modeling

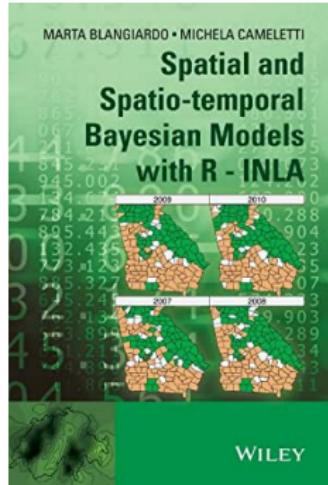
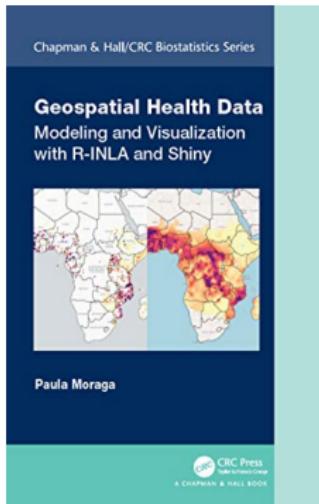
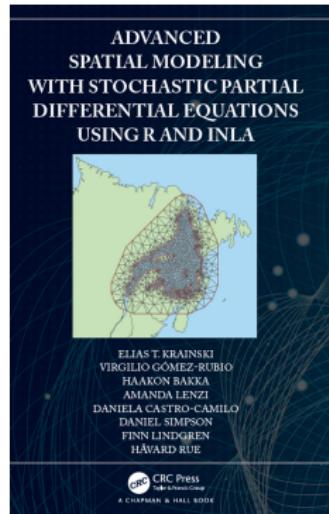
$$\mathbf{Y}_t = \mathbf{M}(\theta_p) \mathbf{Y}_{t-1} + \mathbf{M}_b(\theta_p) \mathbf{Y}_{b,t} + \boldsymbol{\eta}_t$$



## Section 6

Some useful references

# Some useful references



And **Spatio-Temporal Statistics with R** by Christopher K. Wikle, Andrew Zammit-Mangion, and Noel Cressie. Published by Chapman & Hall/CRC.  
<https://spacetimewithr.org/>

## Section 7

### Work in autonomy

## Work in autonomy

- Explore some aspects of spatial statistics that I did not presented here (anisotropy, non-stationarity, EOF, aggregation indices, etc)
- Apply INLA on available dataset (EVHOE or anything else)
- Present some nice modeling/coding stuff based on previous references
  - ➡ 2 hours of work in autonomy and restitution